

RESTRICTED 3-BODY PROBLEM

Due date: 2/1/2018, 9am

Continuing with our exploration of ordinary differential equations, here we address the restricted three-body problem in which a massless particle moves in the gravitational field of two massive objects that orbit their mutual center of mass.

- a) Plot $\psi(r)$ and $\mathbf{F}(\mathbf{r})$ in 2D and 3D.
- b) Determine the location of Lagrangian points and study their stability. In particular, determine the value of $q = M_2/M_1$ for which L_4 and L_5 become unstable.
- c) Using the Runge-Kutta and/or Bulirsch-Stoer integrators, study the orbits of Trojan and Greek asteroids. How sensitive are the orbits to initial conditions? Can you find initial conditions that result in periodic orbits?
- d) Plot the shapes of stars for different values of the equipotential $\Omega(\rho)$:

$$\Omega = \frac{1}{\rho} + q \left[(\rho^2 + 1 - 2\rho\lambda)^{-1/2} - \rho\lambda \right] + \frac{1}{2}\rho^2(1+q)(1-\nu^2).$$

Do not forget that you are solving this in a rotating coordinate system. Hence you need to be careful about the initial conditions.