## EXAMPLES OF NUMERICAL PROBLEMS INVOLVING THE H-R DIAGRAM

The Hertzsprung-Russell diagram in a very fancy form is included on the last page; use it when necessary.

- 1. Let us compare a star with  $2 M_{\odot}$  to the star with  $0.5 M_{\odot}$ .
  - a) Compute the luminosities of both stars. Express your answers in solar luminosity.
  - b) Using the H-R diagram, estimate the surface temperature of both stars. Express the answers in solar temperature.
  - c) Using the luminosity and the temperature, compute the radii of both stars. Express your answers in solar radii.
- 2. Now let us focus on magnitudes.
  - a) How much brighter is a star of magnitude 0 from the star of magnitude 1? 2? 3? 4? 5?
  - b) if the absolute magnitude of a star is larger than the absolute magnitude of the Sun by 5, what is the luminosity of that star, expressed in solar luminosity?
- 3. Next let us focus on the distance modulus.
  - a) Express the distance to the Sun in parsecs.
  - b) What is the absolute magnitude of the Sun if its apparent magnitude is -27?
  - c) If we observe a Sun-like star with the apparent magnitude of 12, how far away is it?
- 4. More fun with the distance modulus.
  - a) Two stars have the same apparent magnitude. One is 3 times farther than the other. What is their luminosity ratio?
  - b) Two stars are at the same distance. One is 3 times brighter than the other. What is their absolute magnitude difference?
  - c) Two stars have the same absolute magnitude. One is 3 times farther than the other. What is their apparent magnitude difference?

- 5. Finally, let us put everything together. Let us again consider our two stars, one with  $2 M_{\odot}$  and the other with  $0.5 M_{\odot}$ .
  - a) Compute again the luminosities of the two stars. Do it again so that you remind yourself that luminosity depends on temperature and size.
  - b) Compute the difference in absolute magnitudes of both stars.
  - c) Compute absolute magnitudes of both stars. You will have to know the absolute magnitude of the Sun for this.
  - d) Compute the distance ratio if both stars appear equally bright.
- 6. One last problem. From the H-R diagram, read off the approximate luminosity of (a) a red supergiant, (b) a red giant, and (c) a red dwarf. Make sure that your reading is done at approximately the same temperature.
  - a) Determine the masses of all three stars; express them in solar mass.
  - b) Determine the radii of all three stars; express them in solar radii.
  - c) If all three stars were at a distance of 10 parsecs, what would their apparent magnitudes be?
  - d) If all three stars were at a distance of our Sun, how much brighter/fainter than the Sun would they be?

HOWEWORK KEX:

$$\Box = (M_0)^{3.5} = 11.3$$

$$\frac{L}{L_0} = \left(\frac{M}{M_0}\right)^{3.5} = 0.09$$

$$\frac{T}{T_0} = 1.5$$

$$\frac{7}{70} = 0.67$$

$$\left(\frac{P_{0}}{P_{0}}\right) = \left(\frac{L}{L_{0}}\right)^{1/2} \left(\frac{T}{L_{0}}\right)^{-2}$$

$$\frac{P}{Po} = 0.67$$

=> 
$$\frac{1}{5}$$
 =  $\frac{2}{5}$  (m, -m<sub>2</sub>)

$$\frac{f_0}{f_1} = 10^{\frac{2}{5}} = 2.512$$
  $\frac{f_0}{f_2} = 10^{\frac{4}{5}} = 6.31$ 

$$\frac{f_0}{f_8} = 10^{\frac{6}{5}} = 15.85$$
  $\frac{f_0}{f_4} = 10^{\frac{8}{5}} = 19.05$ 

$$\frac{f_0}{f_4} = 10\frac{8}{5} = 19.05$$

$$\frac{f_0}{f_5} = 10^2 = 100$$

$$M - M_0 = -\frac{5}{2} \log \frac{L}{L_0}$$

$$= \sum_{L_{\Theta}} = 10^{-\frac{1}{5}(N-N_{\Theta})} = 10^{-\frac{1}{5}S} = 10^{-2} = 0.01$$

$$\frac{1}{D}$$
 | pe  $\frac{1}{D}$  |  $\frac$ 

$$= 50 = 10^{\frac{1}{5}(m-M+5)} = 10^{\frac{1}{5}(12-4.57+5)}$$

$$(4) m_1 = m_2 \qquad m_1 - m_2 = -\frac{5}{2} \log \frac{f_1}{f_2}$$

$$(5) D_1 = 3D_2 \qquad recall that L = f \times 4\pi B^2$$

$$m_1 - m_2 = -\frac{5}{2} \log L_1 \Rightarrow D_2 \qquad recall that L = f \times 4\pi B^2$$

$$m_1 - m_2 = -\frac{5}{2} \log \frac{L_1}{MD_1^2} \frac{MD_2^2}{L_2} = -\frac{5}{2} \log \left(\frac{L_1}{L_2}\right) \frac{D_2}{D_1}$$

$$0 = \log \left(\frac{L_1}{L_2}\right) \left(\frac{1}{3}\right)^2 = 3 \left(\frac{L_1}{L_2}\right) \left(\frac{1}{3}\right)^2 = 1$$

$$\left(\frac{L_1}{L_2}\right) = 9$$

(b) 
$$D_1 = D_2$$
 Since  $D_1 = D_2$ , it does not matter  $M_1 - M_2 = 2$  whether we think of apparent or absolute mags, the difference will  $\frac{f_2}{f_1} = 3$  be the same.

$$m_1 - m_2 = -\frac{5}{2} \log \frac{f_1}{f_2} = -\frac{5}{2} \log \frac{1}{3} = 1.2$$

$$OM_1 = M_2$$
 SHEE  $M_1 = M_2$ , IT FOLLOWS THAT:
$$D_1 = 3D_2$$

$$L_1 = L_2$$

FROM THE INVERSE SQUARE LAW (IDENTICAL TO HA):  $m_1 - m_2 = -\frac{5}{2} \log \left(\frac{L_1}{L_2}\right) \left(\frac{D_2}{D_1}\right)^2 = -\frac{5}{2} \log \left(\frac{1}{3}\right)^2$ 

= 2.38

$$D M_1 - M_2 = -\frac{5}{2} \log \frac{L_1}{L_2} = -\frac{5}{2} \log \frac{11.8 L_2}{0.09 L_2}$$

© 
$$M_1 - M_0 = -\frac{5}{2} \log \frac{L_1}{L_0} = M_1 = M_0 - \frac{5}{2} \log \frac{L_1}{L_0}$$
  
 $M_1 = 2.22$ , and equivalently  $M_2 = 7.46$ 

$$m_1 - m_2 = -\frac{5}{2} \log \frac{f_1}{f_2} = -\frac{5}{2} \log \left(\frac{L_1}{L_2}\right) \left(\frac{D_2}{D_1}\right)$$

$$\frac{L_1}{L_2} = \left(\frac{D_1}{D_2}\right)^2 = \left(\frac{L_1}{L_2}\right)^{1/2} = \left(\frac{|1| \cdot 3}{0.09}\right)^{1/2}$$

$$\frac{MG}{MO} = \left(\frac{LG}{LO}\right)^{13.5} = 3.72$$

$$\frac{MD}{Mo} = \left(\frac{LD}{LO}\right)^{13.5}$$

$$= 0.52$$

$$\frac{\mathcal{L}}{\mathcal{L}_{0}} = \left(\frac{1}{T_{0}}\right)^{4} \left(\frac{2}{R_{0}}\right)^{2}$$

$$\Rightarrow \left(\frac{2}{R_{0}}\right) = \left(\frac{1}{T_{0}}\right)^{1/2} \left(\frac{1}{T_{0}}\right)^{-2}$$

$$\left(\frac{\text{Rag}}{\text{Ro}}\right) = 562 \qquad \left(\frac{\text{Rg}}{\text{Ro}}\right) = 17.8$$

$$\left(\frac{RD}{RO}\right) = 0.56$$

( TO YOU SEE NOW WHY THEY "RE CALLED SUPER-

3 APPARENT MAG @ 10 pc 18 ABSOLUTE MAGNITUDE:

$$M - M_0 = -\frac{5}{2} \log \frac{L_{SG}}{L_0} = -12.5$$

$$M_g - M_o = -\frac{5}{2} \log \frac{L_g}{L_o} = -\frac{5}{2}$$

$$M_0 - M_0 = -\frac{5}{2} \log \frac{L_0}{L_0} = +2.5$$

(3) 
$$m_{sg} - M_{sg} = 5 lg D - 5$$
  
= )  $m_{sg} = M_{sg} + 5 log D - 5$ 

$$m_{g} = M_{0} - \frac{5}{2} l_{g} \frac{L_{g}}{L_{0}} + 5 l_{g} \frac{D}{D} - 5$$

$$= -\frac{31.7}{2} l_{g} \frac{L_{g}}{L_{0}} + 5 l_{g} \frac{D}{D} - 5$$

$$= -\frac{24.2}{2} \frac{L_{g}}{L_{g}} + \frac{5}{2} l_{g} \frac{D}{D} - 5$$

IF YOU SOLVED THIS IN FLUXES, IT'S JUST AN EXTRA STED:

$$m - m_0 = -\frac{5}{2} \log \frac{f}{f_0} = 0$$
 =  $10^{-\frac{2}{5}} (m - m_0)$ 

$$\frac{f_{SG}}{f_0} = 75,852$$
,  $\frac{f_0}{f_0} = 75.8$ ,  $\frac{f_0}{f_0} = 0.08$